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National Superconducting Cyclotron Laboratory Proposal Form - PAC 37

By submitting this proposal, the spokesperson certifies that all collaborators listed have read the Description of Experiment and have agreed to participate in the experiment.

TITLE: Neutron Knockout in Neutron-Rich Ca Isotopes

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Other Experimeters:

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Paul Fallon	LBNL	SR
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Request for primary beam sequence including tuning, test runs, and in-beam calibrations:

Beam 1	Isotope 82Se		Energy 140 MeV/Nucleon	Minimum Intensity 35 pnA	
Sum of Beam Preparation Sum of Beam-On-Target Additional CCF Use Tim Total Time	n Times Times ne	39 hours 142 hours hours 181 hours			
EXPERIMENTAL LOC EXPERIMENTAL EQU	CATION: JIPMENT:	S3 Vault A1900			

GRETINA

S800 spectrograph without scattering chamber

	Setup Time (Days)	Take Down Time (Days)
Access to Experimental Vault:	2	2
Access to Electronic Setup:	2	2
Access to Data Acquisition Computer:	2	2

Date when experiment will be ready to run: 2012-07-01 Dates excluded:

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Detail any modifications needed to the standard configuration of the device used:

Requirements that are outside the current NSCL operating envelope:

Reaction targets at the experimental station:

List any breaks required in the schedule of your experiment:

Non-standard resources:

Other special requirements:

SUMMARY:

We propose to study the neutron-knockout along the Ca isotopic chain, in order to measure the neutron spectroscopic factors in 47,48,49 Ca. Recent calculations by Holt *et al.*, have suggested that the inclusion of 3N forces to describe the structure of neutron-rich Ca isotopes provides an improvement to the predicted properties, such as the magicity of 48 Ca, and the binding energies. These calculations also suggest that in 50 Ca, there should be a reduction of the $f_{7/2}$ neutron occupancy, resulting from a narrowing of the N=28 gap in 50 Ca. We will investigate this prediction by measuring the spectroscopic factors directly, for the 50 Ca ? 49 Ca knockout, and comparing to theoretical expectations. We will also measure knockout from 48 Ca, as a calibration case, as 48 Ca is a good doubly-magic nucleus and the spectroscopic factors are known, and from 49 Ca, to compare the g.s. to g.s. knockout spectroscopic factors with those obtainted in 48 Ca(d,p) 49 Ca reactions.

I. Physics Justification

A central theme in nuclear structure physics is the evolution of shell structure with isospin, and the changes in nucleon single-particle energies as one moves away from stability towards the driplines. The shell model is robust near stability, but has been shown to be modified in nuclei away from the valley of beta stability. Significant effort has been directed, both experimentally and theoretically to understanding the changes in shell structure as a function of neutron excess (see e.g. [So08] and references therein), and a comprehensive picture is developing in many regions of the nuclear chart. However, open questions remain, both specific to the evolution of important nucleon single-particle orbitals, and more generally to the current understanding of the forces required to describe the structure of nuclei.

Of particular interest is the structure of neutron-rich Ca isotopes around N=32, 34. Recent state of the art calculations using the phenomenological interactions GXPF1 and KB3G reproduce well the known level structure up to A=52, but predict different behaviors for 54 Ca. Interestingly, recent theoretical work by Holt et al. [Ho11] has suggested that the Ca isotopes may be in a region of the nuclear chart where 3-body (3N) forces result in important modifications of nuclear structure. NN-only calculations don't reproduce the shell closure at N=28 in the Ca isotopes, and it has been suggested that the modifications required in phenomenological models to reproduce the doubly-magic character of ⁴⁸Ca are largely due to neglected 3N forces. These new calculations no longer support a N=34 shell-closure as is predicted by the GXPF1A phenomenological model, and provide considerable improvements to the theoretical ground state energies of Ca isotopes in the region. It seems natural that mass measurements of these exotic Ca isotopes will provide important information for the theoretical models, an experimental avenue being explored at TRIUMF by Gallant et al [Ga12]. While level structures of ⁴⁹Ca-⁵²Ca are available from deep-inelastic studies [e.g. Br01, Br05, Fo08, Mo11] and knockout reactions [Ga06], we can gain more insight by looking into the single-particle components of the ground and excited-state wavefunctions by direct single neutron knockout reactions. While some data exist for knockout in the Ca isotopes near ⁴⁹Ca in a conference proceeding [Kr08], the picture remains incomplete. Beam intensities are not yet sufficient to directly probe beyond N=32 in the Ca isotopes, but knockout reactions are possible in the lighter isotopes as experimental rates are higher, so that one can reach as far as knockout from ${}^{50}Ca_{30}$.

In Fig. 1, we reproduce the calculated ESPE of Holt *et al.* [Ho11]. An inspection of these calculated levels indicates that the computed single-particle energy-level spacings are quite sensitive to the inclusion of 3N forces in this theoretical framework. As seen in Fig. 1(c), the $2p_{3/2}$ and $1f_{7/2}$ neutron orbitals are nearly degenerate at N=30 when the NN+3N forces are included, while the phenomenological models exhibit a significant gap. In the case that these two neutron orbitals become degenerate and the N=28 gap is quenched in 50 Ca, one would expect mixing between them, and the spectroscopic factors for neutron knockout would be sensitive to the degree of mixing. In other words, with the realistic NN and 3N forces included, while ⁴⁸Ca is doubly-magic, the N=28 gap is quenched already at 50 Ca, and the $1f_{7/2}$ neutron orbital may not be full. (Properly speaking, the erosion of the N=28 gap in the ESPEs seems to be mainly the effect of the realistic NN forces, rather than that of the 3N component.) In this case, the $f_{7/2}$ sum rule would be depleted, which will be evident in the spectroscopic factors for knockout into the state. This qualitative explanation is borne out by the actual calculated spectroscopic factors [Sc12] provided in Fig. 2 for the knockout of ${}^{50}Ca_{gs}$ to ${}^{49}Ca$. Note the large difference expected between the GXPF1 and the realistic force with 3N components. This quenching of the $f_{7/2}$ sum rule may possibly be one of the clearest indications for the role of 3N forces in the Ca isotopes. The results discussed above also show that the $g_{9/2}$ and $f_{5/2}$ levels are very close in energy in ⁵⁰Ca. The recent experiment at Legnaro [Mo11] established a $9/2^+$ state at ~4 MeV that has been interpreted as resulting from a coupling of the 3⁻ phonon in 48 Ca with an unpaired $p_{3/2}$ neutron, where the 3⁻ excitation in 48 Ca is explained as arising from a $s_{1/2}$ or $d_{3/2}$ proton being promoted across the N=20 gap into the $1f_{7/2}$ level. If this is the case, this state should not be populated in the neutron knockout measurement. If however, population of the 4.0 MeV state in ⁴⁹Ca is observed, this would provide first evidence for contributions of the $v1g_{9/2}$ orbital.

II. Goals of the proposed experiment

The goals of the proposed experiment are to measure the low-energy states in the neutron rich 47,48,49 Ca isotopes following neutron knockout, and determine spectroscopic factors for the populated excited states and ground states. This new spectroscopic data will provide unique data to confront the new calculations, describing the $_{20}$ Ca isotopes using realistic NN + 3N forces. It may also identification of possible contributions from the $1g_{9/2}$ neutron orbital which is expected to drop in energy.

III. Experimental Details

We are proposing to use one-neutron knockout reactions ${}^{9}Be({}^{A}Ca, {}^{A-1}Ca+\gamma)X$, for A=48, 49 and 50 and in-beam γ -ray spectroscopy using GRETINA to study the excited states in ${}^{47,48,49}Ca$. We plan to detect the reaction residues in the focal plane of the S800, operated in focus mode, as has been done in the past for these types of reactions. The proposed measurement will involve 3 separate setting of the A1900, centered on ${}^{48,49,50}Ca$, where the primary contaminants in each case will be the V isotope four mass units above the Ca isotope of interest.

Secondary Beam Production and Rates

The secondary beams of 48,49,50 Ca will be produced by projectile fragmentation of a 140 MeV/nucleon ⁸²Se primary beam in a 525 mg/cm² Be target placed at the object position of the A1900 fragment separator. The desired fragments will be separated from other reaction products using a 500 mg/cm² Al wedge at the intermediate dispersive image of the A1900. Simulations using the program LISE++, version 9.3 [Ta04, Ba02], give the production rates for ^{48,49,50}Ca as shown in Table 1, using a 1% momentum acceptance of the A1900. The minimum primary beam current expected for 140 MeV/nucleon ⁸²Se is 35 pnA, giving a lower limit for the production rate of the most exotic nucleus, ⁵⁰Ca, at the A1900 focal plane of ~2460/s. However, based on previous NSCL experiments using a ⁷⁶Ge primary beam in this section of the nuclear chart (NSCL Experiment 05101), LISE++ estimates for yields of rare isotopes were shown to be overestimated in the region by, on average, a factor of 5. Conservatively assuming that the overestimation in rare-isotope yields is similar for the ⁸²Se primary beam, the rate of ⁵⁰Ca expected at the A1900 focal plane is 492/s. The rates for the two secondary beams are given in Table 1, taking into account 50% transmission from the A1900 focal plane to the S3 vault. Identification of the cocktail beams impinging on the S800 target position will be achieved using the difference in time-of-flight between the A1900 extended focal plane and the S800 object scintillators. This method has been used successfully in previous S800 knockout experiments.

Experiment Set-Up

This measurement will use the combination of GRETINA and the S800 spectrograph for coincident particle- γ spectroscopy. The seven GRETINA modules will be used at forward angles, with four detectors in the first ring at $\theta = 58^{\circ}$, and the other three in a close configuration in the second ring at $\theta = 90^{\circ}$. The high resolution of GRETINA as compared to CAESAR will

allow us to resolve closely-spaced gamma-ray transitions near 4 MeV, while the efficiency gain as compared to SeGA makes the measurement feasible within a reasonable length of beam time. At the target position of GRETINA, a 400 mg/cm² ⁹Be target will induce neutron-knockout reactions. Reaction residues will be detected and identified in the S800 focal plane detectors, with additional information coming from beam-line timing detectors. Reconstruction of flight trajectories through the S800 and momentum distributions will be achieved using positions and angles in the focal plane as measured using the position-sensitive CRDC detectors, and inverse maps for the spectrograph.

Reaction Rates and Beam Time Request

We propose to first measure the neutron knockout for ⁴⁸Ca to ⁴⁷Ca. ⁴⁸Ca is a good doubly-magic nucleus, and knockout into ⁴⁷Ca will thus serve as an excellent calibration measurement, as the states in ⁴⁷Ca are also well known, as are the spectroscopic factors. With the high rate of ⁴⁸Ca, only 4 hours of beam on target time is requested for this calibration measurement. We will then measure the knockout for ⁴⁹Ca to ⁴⁸Ca -- the ⁴⁸Ca(d,p)⁴⁹Ca reaction [Uo94] has been studied in detail, and we will be able to confirm the results of that work by studying the complementary knockout. Our results for the g.s. to g.s. spectroscopic factors should be consistent with those extracted in the (d,p) measurement. With a secondary beam rate of 1100 pps at the S800 target position, we request 24 hours to perform a detailed measurement of the ${}^{49}Ca \rightarrow {}^{48}Ca$ knockout. For the target case of ${}^{50}Ca \rightarrow {}^{49}Ca$, to make a detailed momentum distribution measurement and extract spectroscopic factors for a given state, we require 1000 particle-gamma coincidence events. Assuming a nominal 10 mb single-particle cross-section, estimates for the expected gamma-ray rates are presented in Table 2 for the knockout of ⁵⁰Ca to ⁴⁹Ca. We request 100 hours of beam time to study this knockout reaction. With this beam time, we will be able to make a high-statistics measurement for the 3.36 MeV state, which will give us information regarding the occupancy of the $1f_{7/2}$ orbital, which is expected to be reduced if 3N forces close the N=28 gap as predicted. In addition, we expect to obtain a few hundred counts in the 3.99MeV distribution, sufficient to determine the spectroscopic factor for $1f_{5/2}$ in the ⁵⁰Ca ground state, as this will be much less than the 1 assumed in the gamma-ray rate estimate of Table 2. With 100 hours of beam time, and the efficiency of GRETINA, we will also be sensitive to any possible population of the 4.0 MeV $9/2^+$ state, if there is any contribution of $v1g_{9/2}$ in this excited state configuration, and the ⁵⁰Ca ground state.

IV. Supplemental Information (Figures, Tables, References, etc., including one figure that depicts the layout of the experimental apparatus)

References

[Ba02] D. Bazin et al. Nucl. Instrum. Methods Phys. Res. A 482, 307 (2002).

[Br01] R. Broda, Acta Phys. Pol. B 32, 2577 (2001).

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[Cr12] H. Crawford et al., to be published. Based on R. Clark et al., NSCL E09032.

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[Ga06] A. Gade et al., Phys. Rev. C 74, 021302R (2006).

[Ga12] A. Gallant et al., arXiv:1204.1987 (2012).

[Ha99] M. Hannawald et al., Phys. Rev. Lett. 82, 1391 (1999).

[Ho11] J. Holt et al., arXiv:1009.5984v2 (2011).

[Kr08] R. Krucken, *AIP Conf. Proc.*, "Nuclear Physics and Astrophysics: From Stable Beams to Exotic Nuclei" **1072**, 52 (2008).

[Mo11] D. Montanari et al., Phys. Lett. B 697, 288 (2011).

[Ro11] W. Rother et al., Phys. Rev. Lett. 106, 022502 (2011).

[Sc12] A. Schwenk et al., private communication.

[So08] O. Sorlin et al., Prog. Part. Nucl. Phys. 61, 602 (2008).

[Ta04] O. Tarasov and D. Bazin. Nucl. Phys. A 746, 211 (2004).

[Uo94] Y. Uozumi et al., Nucl. Phys. A 576, 123 (1994).



Figure 1: Neutron single-particle energy levels relative to those in 40 Ca, as calculated using (a) phenomenological forces, fit to data; (b) NN-only calculations; (c) NN + 3N calculations, and (d) NN + 3N calculations in an expanded model space. Taken from Reference [Ho11].





Figure 2: Spectroscopic factors calculated using the single-particle energy levels based on phenomenological forces (GXPF1), NN+3N forces restricted to the *fp*-shell, and including the $1g_{9/2}$ orbital. The reduction of the *N*=28 shell gap with the inclusion of 3N forces is expected to result in the depletion of the $1f_{7/2}$ sum rule in ⁵⁰Ca, and a reduced spectroscopic factor for knock-out into ⁴⁹Ca.



Figure 3: Level structure of ⁴⁹Ca from [Mo11].

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NSCL PAC 37 - 2. Status of Previous Experiments

Tuble 1. Secondary beam production mornation for a secondary beams				
Isotope	⁴⁸ Ca	⁴⁹ Ca	⁵⁰ Ca	
LISE++ Rate with 1% dp/p [pps/pnA]	1290	316	70.4	
LISE++ Purity	92%	79%	77%	
Expected Rates at A1900 Focal Plane [pps]	9030	2212	492	
Rates at S800 Target [pps]	4515	1106	246	

Table 1: Secondary beam production information for 4 secondary beams

Table 2: Knockout for ${}^{50}Ca \rightarrow {}^{49}Ca$, assuming the level structure of [Mo11].

E [keV]	Configuration	Expected Occupancy (C ² S)	Rate [1/hour]	Eγ (keV)	GRETINA Efficiency	Gamma Rate [1/hour]
		⁴⁸ Ca	$a \rightarrow {}^{47}Ca$			
0	$1f_{7/2}^{1}$	8	34784			
2569	$2d_{3/2}^{1}$	4	17392	2569	6%	1044
2591	$2s_{1/2}^{1}$	2	8696	2591	6%	522
	${}^{50}Ca \rightarrow {}^{49}Ca$					
0	$1p_{3/2}^{1}$	1	237			
2023.2	$1p_{1/2}^{1}$	1	237	2023.2	6.5%	15.4
3357	$[f_{7/2}^{-1}p_{3/2}^{2}]$	4 / 2 [Sc12]	948	3357	5%	47.4
3991	$1 f_{5/2}^{1}$	<1	237	3991	4%	9.5 (if $C^2S = 1$)
4017	$3^{-} \otimes 1p_{3/2}^{-1}$					



Figure 4: GRETINA at the S800

Status of Previous Experiments

Results from, or status of analysis of, previous experiments at the CCF listed by experiment number. Please indicate publications, invited talks, Ph.D.s awarded, Master's degrees awarded, undergraduate theses completed.

E09020: Coulomb Excitation of Neutron-rich Fe and Cr (March, 2011)

- Analysis is nearly completed, publication is in preparation
- Results presented in oral contribution at Rutherford Centennial and Gamma11 international conferences

NSCL PAC 37 – 3. Educational Impact

Educational Impact of Proposed Experiment

If the experiment will be part of a thesis project, please include the total number of years the student has been in graduate school, what other experiments the student has participated in at the NSCL and elsewhere (explicitly identify the experiments done as part of thesis work), and what part the proposed measurement plays in the complete thesis project.

Students local to MSU/NSCL will be invited to participate in the set-up and execution of the experiment. The experiment will not be a part of any thesis work.

NSCL PAC 37 – 4. Safety Information

Safety Information Worksheet

It is an important goal of the NSCL that users perform their experiments safely, as emphasized in the <u>Director's Safety Statement</u>. Your proposal will be reviewed for safety issues by committees at the NSCL and MSU who will provide reviews to the PAC and to you. If your experiment is approved, a more detailed safety review will be required prior to scheduling and you will need to designate a <u>Safety Representative</u> for your experiment.

SAFETY CONTACT FOR THIS PROPOSAL: Dirk Weisshaar

HAZARD ASSESSMENTS (CHECK ALL ITEMS THAT MAY APPLY TO YOUR EXPERIMENT):

X	Radioactive sources required for checks or calibrations. Transport or send radioactive materials to or from the NSCL. Transport or send— to or from the NSCL—chemicals or materials that
	may be considered hazardous or toxic.
	Generate or dispose of chemicals or materials that may be considered
hazardous or toxic.	
	Mixed Waste (RCRA) will be generated and/or will need disposal.
	Flammable compressed gases needed.
	High-Voltage equipment (Non-standard equipment with > 30 Volts).
	User-supplied pressure or vacuum vessels, gas detectors.
	Non-ionizing radiation sources (microwave, class III or IV lasers, etc.).
	Biohazardous materials.
	Lifting or manipulating heavy equipment (>500 lbs)

PLEASE PROVIDE BRIEF DETAIL ABOUT EACH CHECKED ITEM.

Standard gamma-ray sources will be used for energy and efficiency calibration of GRETINA.

NSCL PAC 37 – 5. Spectrograph Worksheet

Spectrograph Worksheet for S800 Spectrograph or Sweeper Magnet

The NSCL web site contains detailed technical information and service level descriptions about the <u>S800 Spectrograph</u> (<u>Service Level Description</u>) and the <u>Sweeper Magnet</u> (<u>Service Level Description</u>).

1. Timing detectors

a) Is a plastic timing scintillator required (at the object of the S800 or in front of the sweeper magnet)?

[] No

[X] Yes

- i. What is the desired thickness? [X] 125 µm [] 1 mm [] other _____
- ii. What maximum rate is expected on this scintillator? _____ Hz

b) Do you plan to use a different type of timing detector (at the object of the S800 or in front of the sweeper magnet)?

[X] No

[] Yes

If "Yes," please give details.

2. Tracking detectors

Tracking detectors for incoming beam are available for Z>10. Performance limitations are to be expected at rates exceeding 200 kHz.

Are tracking detectors needed?

[] No

[X] Yes

3. Focal-plane rates

a) What detectors are planned to be used?

Standard FP detectors: CRDCs, Ion Chamber, Plastic Scintillator and CsI Hodoscope

b) What is the maximum rate expected in the focal-plane detection system? 1000 Hz

4. For S800 experiments only: Optics mode and rigidities:

- a) Which optics mode is needed?
 - [] Dispersion matched [X] Focused [] Other _
- b) What are the maximum and minimum rigidities planned to be used for the analysis beam line?

3.2 Tm minimum, 3.5 Tm maximum

c) What are the maximum and minimum rigidity planned to be used for the spectrograph?

2.5 Tm minimum, 2.9 Tm maximum

d) The maximum particle rate in the focal plane is 6 kHz when the CRDC detectors are being used. What is the maximum total particle rate expected in the S800 focal plane? < 1000 Hz</p> NSCL PAC 37 - 6. Beam Request Worksheet Instructions

Beam Request Worksheet Instructions

Please use a separate worksheet for each distinct beam-on-target requested for the experiment. Do not forget to include any beams needed for calibration or testing. This form does not apply for experiments based in the A1900. Note the following:

- (a) **Beam Preparation Time** is the time required by the NSCL for beam development and beam delivery. This time is calculated as per item 4. of the Notes for PAC 37 in the Call for Proposals. This time is not part of the time available for performing the experiment.
- (b) **Beam-On-Target Time** is the time that the beam is needed by experimenters for the purpose of performing the experiment, including such activities as experimental device tuning (for both supported and non-supported devices), debugging the experimental setup, calibrations, and test runs.
- (c) The experimental device tuning time (XDT) for a supported device is calculated as per item 5. of the Notes for PAC 37 in the Call for Proposals. For a non-supported device, the contact person for the device can help in making the estimate. In general, XDT is needed only once per experiment but there are exceptions, e.g. a change of optics for the S800 will require a new XDT. When in doubt, please consult the appropriate contact person.
- (d) A **primary beam** can be delivered as an on-target beam for the experiment either at the full beam energy or at a reduced energy by passing it through a degrader of appropriate thickness. The process of reducing the beam energy using a degrader necessarily reduces the quality of the beam. Please use a separate worksheet for each energy request from a single primary beam.
- (e) Report the Beam-On-Target **rate** in units of particles per second per particle-nanoampere (pps/pnA) for secondary beams or in units of particle-nanoampere (pnA) for primary or degraded primary beams.
- (f) More information about **momentum correction** and **timing start signal** rate limits are given in the <u>A1900 service level description</u>.
- (g) For rare-isotope beam experiments, an electronic copy of the LISE++ files used to estimate the rare-isotope beam intensity must be e-mailed to the <u>A1900 Device Contact</u>.

NSCL PAC 37 – 6. Beam Request Worksheet

Beam Request Worksheet <u>1</u> of <u>3</u>.

(Please number the sheets and use a separate sheet for each distinct beam-on-target requested.)

		Beam Preparation Time	Beam- On-Target Time
Primary Beam (from beam list)			
Isotope 82Se			
Energy 140	MeV/nucleon		
Minimum intensity35	particle-nanoampere (pnA)		
Tuning time (12 hrs; 0 hrs if the	beam is already listed in an earlier worksheet):	12 hrs	
Beam-On-Target			
Isotope 48Ca			
Energy 81	MeV/nucleon		
Rate at A1900 focal plane 1290	pps/pnA (secondary beam) or pnA (primary	beam)	
Total A1900 momentum acceptance 1	% (e.g. 1%, not $\pm 0.5\%$)		
Purity at A1900 focal plane 92	%		
Is a plastic timing scintillator required at the A1900 foc [] No [X] Yes What is the desired thickness? [What is the maximum rate expected f	xal plane for providing a timing start signal? X] 125 μm; [] 1000 μm for this setting? _20000_Hz (1 MHz max)		
Is event-by-event momentum correction from position : [X] No [] Yes	measured at the A1900 Image 2 position required	1?	
Which detector should be used? [What is the maximum rate expected f] Scintillator; [] PPACs for this setting?Hz (1 MHz max)		
Delivery time per table (or 0 hrs	for primary/degraded primary beam):	6 hrs	
Tuning time to vault:		3 hrs	
Total beam preparation time f	or this beam:	21 hrs	
Experimental device tuning time S800 [X]; SeGA []; Sweeper On-target time excluding device	e [see note (c) above]: []; Other [X] GRETINA tuning:		6 hrs 4 hrs
Total on-target time for this be	eam:		10 hrs

NSCL PAC 29 Beam Request Worksheet

Beam Request Worksheet <u>2</u> of <u>3</u>.

(Please number the sheets and use a separate sheet for each distinct beam-on-target requested.)

		Beam Preparation Time	Beam- On-Target Time
Primary Beam (from <u>beam list</u>)			
Isotope 82Se			
Energy <u>140</u>	MeV/nucleon		
Minimum intensity 35	particle-nanoampere (pnA)		
Tuning time (12 hrs; 0 hrs if the bea	am is already listed in an earlier worksheet):	0 hrs	
Beam-On-Target			
Isotope <u>49Ca</u>	NA 37/ 1		
$\frac{\text{Energy}}{\text{Rate at } 41900 \text{ focal plane}} = \frac{316}{316}$	Me $\sqrt{\text{nucleon}}$	eam)	
Total A1900 momentum acceptance 1	$\%$ (e.g. 1%, not $\pm 0.5\%$)	(calli)	
Purity at A1900 focal plane 75	%		
Is a plastic timing scintillator required at the A1900 focal []] No [X] Yes What is the desired thickness? [X] What is the maximum rate expected for	plane for providing a timing start signal? 125 µm; [] 1000 µm this setting? _20000_Hz (1 MHz max)		
Is event-by-event momentum correction from position me [X] No	asured at the A1900 Image 2 position required	?	
[] Tes Which detector should be used? [] What is the maximum rate expected for	Scintillator; [] PPACs this setting?Hz (1 MHz max)		
Delivery time per table (or 0 hrs for	r primary/degraded primary beam):	6 hrs	
Tuning time to vault:		3 hrs	
Total beam preparation time for	this beam:	9 hrs	
Experimental device tuning time [s S800 [X]; SeGA []; Sweeper [On-target time excluding device tur	ee note (c) above]:]; Other [] ning:		4 hrs 24 hrs
Total on-target time for this bean	n:		28 hrs

NSCL PAC 29 Beam Request Worksheet

Beam Request Worksheet <u>3</u> of <u>3</u>.

(Please number the sheets and use a separate sheet for each distinct beam-on-target requested.)

	Beam Beam- Preparation On-Target Time Time
Primary Beam (from <u>beam list</u>) Isotope <u>82Se</u> Energy <u>140</u> MeV/nucleon Minimum intensity <u>35</u> particle-nanoampere (pnA)	· · · · · · · · · · · · · · · · · · ·
Tuning time (12 hrs; 0 hrs if the beam is already listed in an ear	rlier worksheet): 0 hrs
Beam-On-TargetIsotope50CaEnergy82.5MeV/nucleonRate at A1900 focal plane70pps/pnA (secondary beam)Total A1900 momentum acceptance1% (e.g. 1%, not ±0.5%)Purity at A1900 focal plane75%	or pnA (primary beam)
Is a plastic timing scintillator required at the A1900 focal plane for providing a timing []] No [X] Yes What is the desired thickness? [X] 125 µm; [] 1000 µm What is the maximum rate expected for this setting? _3000_Hz (1)	; start signal? MHz max)
Is event-by-event momentum correction from position measured at the A1900 Image [X] No [] Yes Which detector should be used? [] Scintillator; [] PPACs What is the maximum rate expected for this setting? Hz	2 position required? z (1 MHz max)
Delivery time per table (or 0 hrs for primary/degraded primary	beam): 6 hrs
Tuning time to vault:	3 hrs
Total beam preparation time for this beam:	9 hrs
Experimental device tuning time [see note (c) above]: S800 [X]; SeGA []; Sweeper []; Other [] On-target time excluding device tuning:	4 hrs 100 hrs
Total on-target time for this beam:	104 hrs